

Data transfer from ALMA to North America



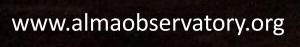
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Observatory









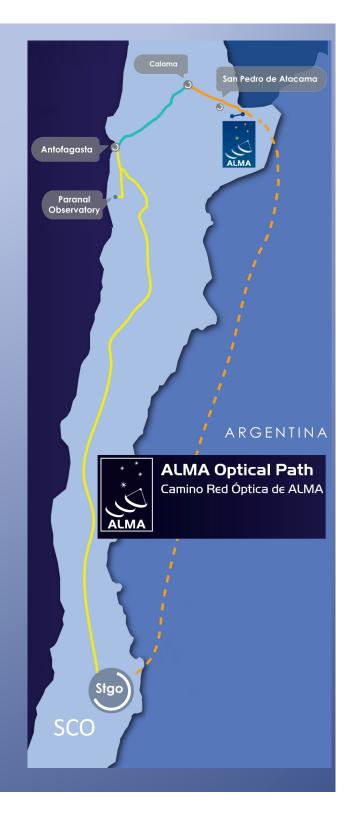


ALMA telescope

- Largest mm/submm telescope ever built.
- Interferometer combines signals from multiple antennas to form an image.
- All 66 antennas delivered, all at high site (except for maintenance).
- Multinational project with many partners, three ALMA Regional Centers (ARCs): US, EU and EA
- Operated "space mission" style, with pipeline data processing and a science archive at each ARC allowing data reuse.
- First PI projects released to public from the ARCs January 2013
- Cycle 5 observations began in October 2017.

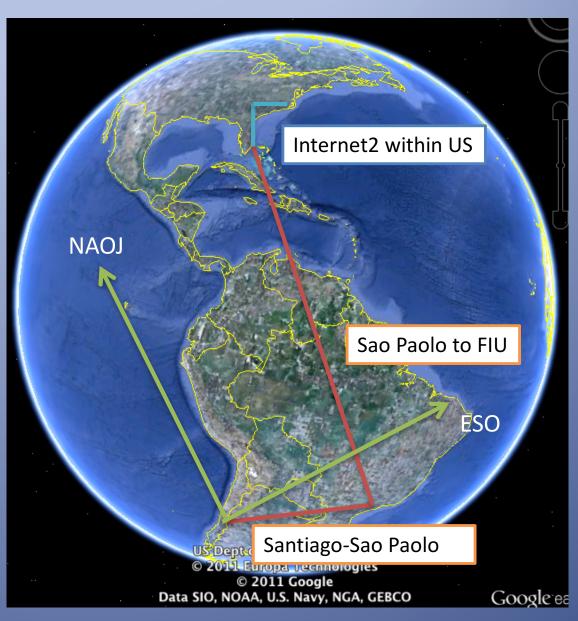
Data Transfer within Chile

- AOS to Santiago 2.5Gb/s (fiber to Calama, commercial fiber Calama to Antofagasta, EVALSO/REUNA from Antofagasta to Santiago.
 - Redundant fiber loop via Argentina planned.
- Santiago to ARCs: individual ARC contracts with REUNA and NRENs
- Data processing to produce Level 2/3 products shared between Santiago and the ALMA regional centers.
 - Pipeline is run at 4 locations worldwide, including Santiago.
 - Data packages ingested into the archive in Santiago.
 - Pipeline products ~ same size as raw data.
 - A significant amount of data (~100TB/yr) is thus being uploaded from the NA ARC to JAO. Bi-directional speed is thus important.
- Long-term plan is that all data processing will take place in Santiago.

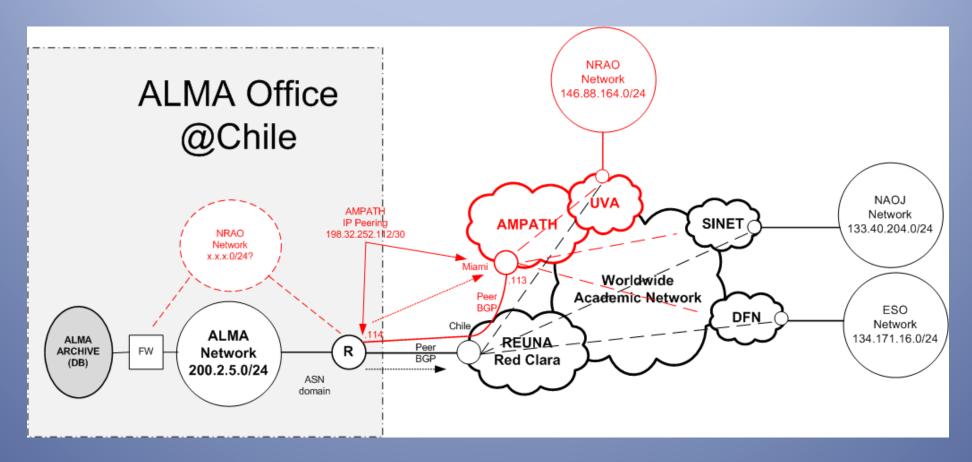


Data transfer – Chile to NA

- Joint AURA-AUI agreement for 100Mb/s committed (burstable to capacity) of AURA's 622Mb/s link to Chile through Sao Paolo and Miami (FIU/AmLight) to the US research network backbone (NREN).
- MOU signed between AUI/REUNA for local link to SCO.
- Link from NRAO to Internet2 through UVa is 10Gb/s
- Typical rate obtained during peak data transfer periods is 2-300Mb/s, with bursts up to 600Mb/s.
 - 90% is bulk data with low QoS.
 - Remainder is database sync and telepresence
- Currently working on establishing network monitoring, and improving our understanding of how the link performs in high load conditions (~1TB/day).



Paths from ALMA



Note: NRAO will be abandoning 146.88.x.x IP space in the next month

ALMA Science data rate evolution

- ALMA Cycle 0 completed (Oct 2011-Jan2013)
 - 16-24/50 antennas used (data rate proportional to square of antenna number)
 - ~5-10% of array time for science
 - Total data volume was about 20TB
- ALMA Cycle 1 complete (~Aug 2013-Jun2014)
 - 32-40/50 antennas, plus 7/12 compact array
 - ~10% of array time for science
 - 40TB over 1yr (ALMA archive hit 50TB in March 2014)
- ALMA Cycle 2 complete (June 2014-Sept 2015)
 - ~34 main array antennas, 10 compact array
 - ~15% of array time for science (but some carryover from Cycle 1)
 - 70TB in a 17 month Cycle.
- ALMA Cycle 3 complete (Oct 2015-Sept 2016).
 - 36 main array antennas, 10 compact array
 - ~25% of array time for science
 - Total of 140TB, mostly raw data (manual imaging and imaging pipeline products only ~20%).
 - Data volume artificially high as two data streams are kept with different corrections.
- ALMA Cycle 4 complete (Oct 2016-Sept 2017)
 - 40 main array antennas, 10 compact array
 - ~33% of array time for science
 - Total of 210 TB, mix of raw and pipeline image products
 - Data volume artificially high as two data streams are kept with different corrections.

Cycle 5

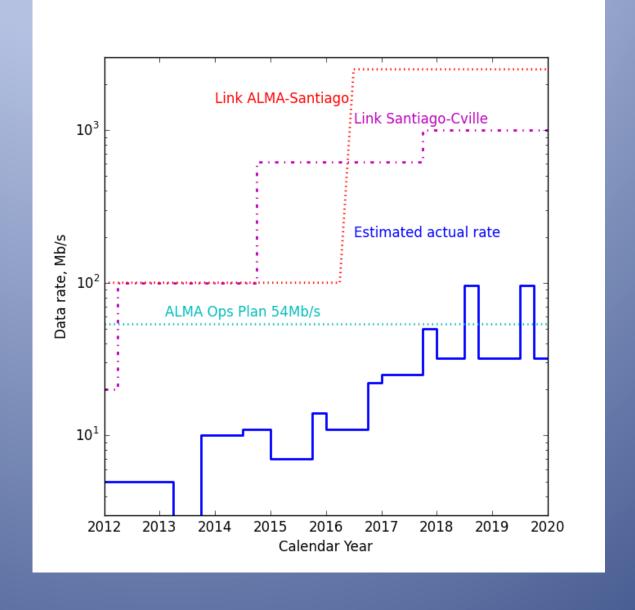
- Still taking data in 2 streams (WVR corrected and uncorrected)
- ~45% of array time for science
- More antennas (43 compared to 40).
- Total volume will be ~360TB, including image products (which will constitute about 50% of the total data volume).

Future Cycles

- Raw data rates will continue to increase as we transition into full operations. Expect Full Science cycles (Cycle 6 onwards) to have mean data rates ~100Mb/s during observations, but could be ~50% higher. "Duty cycle" of observations will also increase as testing and maintenance procedures improve.
- Best guess estimate (including product size mitigation) is around 400TB (200TB raw, 200TB products).
- Important to note that data rates vary through the configuration cycle. When long baseline configurations are scheduled the data rate goes up for two reasons:
 - Data sampling needs to be faster to prevent beam smearing at the field edges.
 - The data products, which are also mirrored from Santiago, also increase in size, to become larger than the raw data in the largest configurations.
 - So far, long baseline campaigns have tended to have low observing efficiencies, however this may change.

Current data rate projections

- Assumes no imposed limit on data rate (cyan line is current Operations Plan rate).
- Blue line is for data generation
- Data transmission is per ARC



Summary

- Ramp-up of the ALMA data rate has been slower than anticipated, allowing us to stay ahead of the curve.
- 2-stream data collection in Cycles 3-5 is artificially boosting the data rate, we are assuming this will no longer be the case in Cycle 6 (Oct 2018 onwards).
- Bidirectional data flow will continue to be needed to support data processing at the ARCs.
- Still learning how the network performs when transferring ~1TB/day in multiple parallel streams.
- Would like to establish a link with 1Gb/s available bandwidth (out of a 10Gb/s pipe) within the next 1-2 years to improve our transfer speed to and from Chile for bulk reprocessing, and to help with occasional large data and metadata transports (e.g. a DB export).
- Most new developments (e.g. correlator improvements) on <10yr timescales can probably be accommodated without increasing the data rate by more than a factor ~2-4.